# Decay-Scheme of <sup>130</sup>Cd

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With the technical improvements and results described in the preceding annual report about "First  $\gamma\text{--Spectroscopic}$  Study of the N=82 r-Process "Waiting-Point" Nucleus  $^{130}\text{Cd}$ ", we were able to collect  $\beta\gamma\text{--}$ coincidence data and to determine the  $Q_\beta\text{--value}$  for  $^{130}\text{Cd}$ . This result completed our earlier investigations and allowed to propose a first  $\beta\text{--decay}$  scheme of  $^{130}\text{Cd}$ , that can be compared to shell-model predictions.

#### Decay scheme

The most important nuclear physics quantities in this context are the excitation energy of the lowest 1 + 2QP state and its log ft value for Gamow-Teller (GT) feeding. The input data needed to calculate the log ft are the GT-decay energy, which is deduced from the difference between the  $Q_{\beta}$ -value and the E(1<sup>+</sup>), the  $\beta$ -decay half-life of 162 ms and the  $\beta$ -feeding of 69%. The predictions for the Q $_{\beta}$ -value of 130°Cd lie between 8.93 MeV from mass models with shellquenching (e.g. HFB-SkP and ETFSI-Q) and 7.43 MeV from the unquenched FRDM, while the Audi and Wapstra mass evaluation [1] leads to a value of 8.5 MeV. Our experimental value of 8.32 MeV [2] was deduced from a simplified Fermi-Kurie plot (see Fig.1) and confirms those high values from quenched models. This  $Q_{\beta}$  value is needed in order to obtain a physically consistent picture of  $^{130}$ Cd β-decay.

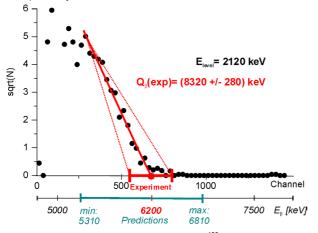


Fig. 1: Simplified Fermi-Kurie plot for  $^{130}\text{Cd}.$  Shown is the summation over all five transitions, that come from the 1 $^{+}$  level at 2120 keV. From its x-axis intercept, a  $\beta\text{-endpoint}$  energy of 6200 keV was derived.

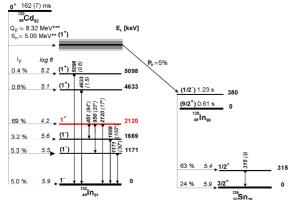


Fig. 2: Decay scheme of <sup>130</sup>Cd. The S<sub>n</sub>-value is taken from [1].

With these data, a log ft value of 4.2 for the  $0^+ \rightarrow 1^+$  GT-transition was deduced. The  $\beta$ -feeding of the ground-state suggests a first-forbidden (ff)-transition, while the log ft's of the levels at 1669 keV and 1171 keV lie between the usual values for GT- and ff-decays.

### **OXBASH** model predictions

Fig.3 shows a comparison of the experimental level scheme of  $^{130}$ In with two calculations of the OXBASH model [3,4]. Prediction (A) was made before knowing our results, and the position of the  $[\pi g_{9/2} \otimes \nu g_{7/2}]$  2QP 1<sup>+</sup>-state was calculated to 1382 keV. This theoretical E(1<sup>+</sup>) was considerably lower than our experimental value of 2120 keV. In the recalculation labelled (B), the 1<sup>+</sup> level could be lifted to 1874 keV by weakening the interaction of the  $[\pi g_{9/2} \otimes \nu g_{7/2}]$  two-body matrix element. Apart from the above state, the three low-lying 1<sup>-</sup> levels are also of interest, because the may be populated by ff-transitions. The two high-lying 1<sup>+</sup>-states of 4QP nature at 4633 keV and 5098 keV are beyond the present model calculations, since they require corebreaking.

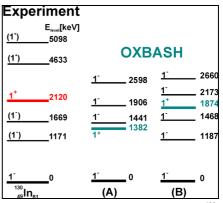


Fig. 3: Comparison of low-lying  $1^+$  and  $1^-$  levels in  $^{130}$ In with two calculations of the OXBASH shell model. Marked in red and blue are the positions of the lowest  $1^+$  level.

## Astrophysical implications

The high experimental  $Q_{\beta}\text{-value}$  of 8.32 MeV is another evidence for a "quenching" of the N=82 shell below  $^{132}\text{Sn}$  [5]. As a consequence of the weakening of the  $\pi\nu\text{-monopole}$  interaction, the half-lives for GT-decay of the so far unknown N=82 waiting-point nuclei  $^{128}\text{Pd},$   $^{127}\text{Rh},$   $^{126}\text{Ru}$  and  $^{125}\text{Tc}$  will become <code>longer</code> than predicted by recent shell models. This leads to a more reliable description of the build-up of the A≈130 N<sub>r,©</sub> peak in r-process abundance calculations. Hence, a better understanding of the r-process matter flow through this "bottle-neck" region is derived, which also determines to a large extent the total duration of a classical r-process.

### References

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