

How doubly-magic is the nucleus ${}^{78}_{28}\text{Ni}_{50}$?

B. Pfeiffer^{1,2} and K.-L. Kratz^{1,2}

¹Institut für Kernchemie, Universität Mainz, Germany; ²HGF VISTARS

Doubly-magic nuclei play an important role in nuclear structure theory as testbeds for shell model calculations. Data on neutron-rich nuclei in the regions of doubly-magic nuclei such as ${}^{78}_{28}\text{Ni}_{50}$ and ${}^{132}_{50}\text{Sn}_{82}$ have a decisive influence on nucleosynthesis calculations. These “longer-lived” waiting-point nuclei determine the duration of the r-process and the matter flow through the abundance maxima at the magic neutron numbers [1]. Remaining deficiencies prior to the abundance maxima in r-process calculations have been interpreted as signatures of new nuclear structure effects near the neutron drip-line, for example overestimation of the shell strength far from stability in global mass models such as FRDM and ETFSI-1. A

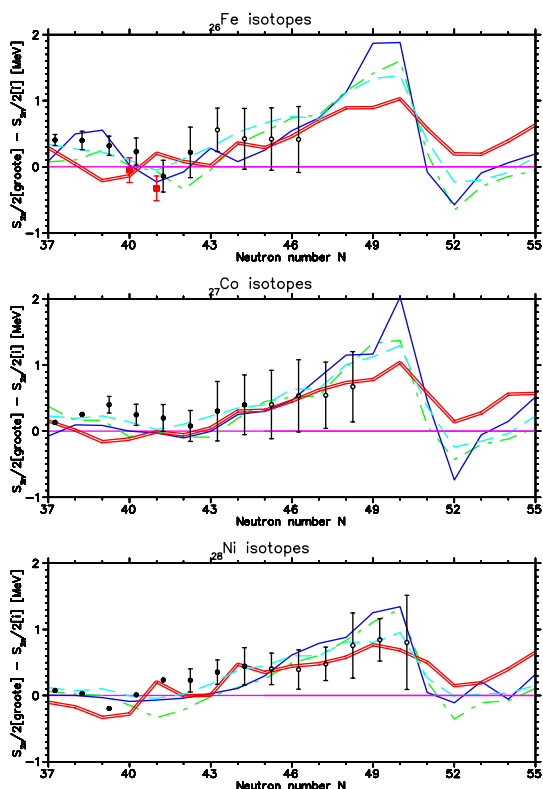


Figure 1: Deviations of experimental [black dots: 2003 mass evaluation [7], red squares: FRS/ESR measurements [8]] and theoretical S_{2n} values from the smoothly varying Groote mass formula [2] across the shell gap at $N=50$ are shown for ${}_{26}\text{Fe}$ (upper part), ${}_{27}\text{Co}$ (middle part) and ${}_{28}\text{Ni}$ (lower part) isotopes. [Theoretical masses: Groote: magenta, FRDM: red [3], ETFSI-1: cyan [4], HFB-2: green [5], HFB-8: blue [6]]

weakening (“quenching”) of spherical shells with increasing isospin, resulting in a gradual setting in of collectivity, has been predicted by recent HFB calculations, and is well established for the lower neutron-magic numbers $N=20$ and $N=28$.

In this context, it is of interest to determine the mutual influence of the proton and neutron magic numbers far from stability. The shell strength can be derived from the dif-

ferences of the two-neutron separation energies S_{2n} prior and behind a magic neutron number. Fig. 1 displays differences between experimental and theoretical S_{2n} values relative to the smoothly varying mass formula of Groote et al. [2] for Fe, Co and Ni isotopes.

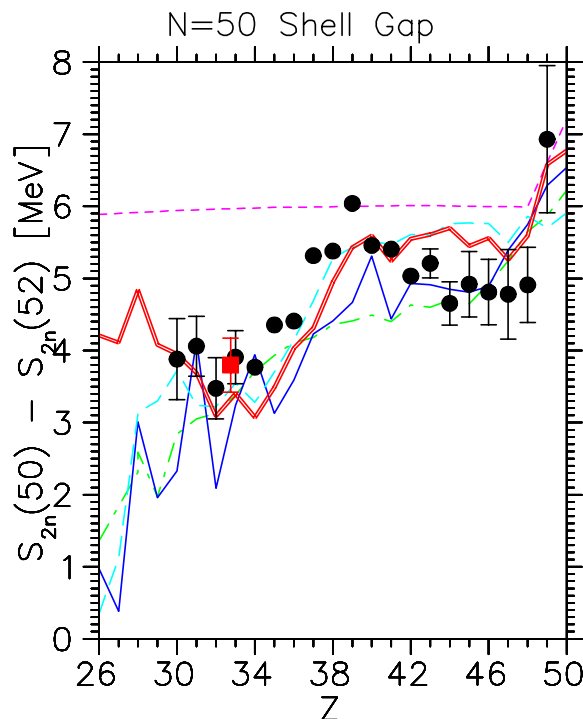


Figure 2: The $N=50$ shell gap as a function of Z . Experimental and extrapolated values from the 2003 mass evaluation [7] (black circles) and experimental values from direct mass measurements at ESR/FSR [red square] are compared to theoretical mass models. [Same colour coding for the mass models as in Fig. 1.]

Fig. 2 shows the $N=50$ shell gap over a wide Z range. Most global mass models predict a local maximum for doubly-magic ${}^{78}\text{Ni}$. So far, below $Z=30$ no experimental masses have been determined, so that this prediction cannot be verified. With upgraded U-beams, direct mass measurements at the FRS-ESR of GSI will in future extend the range of experimental masses in this region to more neutron-rich isotopes.

References

- [1] K.-L. Kratz et al., *Ap. J.* **403**, 216 (1993).
- [2] H. von Groote et al., *ADNDT* **17**, 418 (1976).
- [3] P. Möller et al., *ADNDT* **66**, 131 (1997).
- [4] Y. Aboussir et al., *ADNDT* **61**, 127 (1995).
- [5] S. Goriely et al., *ADNDT* **77**, 311 (2001).
- [6] M. Samyn et al., *Phys. Rev.* **70**, 044309 (2004).
- [7] G. Audi et al., *Nucl. Phys.* **A729**, 3 (2003).
- [8] M. Matos, PhD thesis, Giessen (2004).